

Introduction

Neutrinos are extremely light, electrically neutral leptons which only interact through gravity and the weak force. Due to their lack of charge, neutrinos can pass through most forms of matter. Neutrinos form from nuclear reactions in supernovas and stars (Ahmad et. al, 2001). Solar neutrinos are those which originate from the nuclear reactions within the Sun, and every second, approximately 65 billion solar neutrinos pass through a square centimeter. Most current forms of neutrino detection take place deep underground in large tanks of water, surrounded by photomultiplier tubes (PMTs), to escape background noise from showers produced by cosmic rays and radiation.

Among the particles produced in cosmic ray showers are muons, a type of charged lepton. Muons are produced with various momenta, and this determines their probability of reaching sea level. Muons with momenta around 100 GeV have a 97% probability of surviving to sea level (Allkoißf et. al). Their presence compromises the detection of neutrinos (Tanaka et. al, 2007). Muons are used as a metric for shielding against secondary cosmic rays.

One distinction between neutrinos and particles which compose cosmic rays and radiation is that they are electrically neutral. By exploiting this characteristic, it may be possible to construct neutrino detectors that are not deep underground. A prototype electromagnetic shielding apparatus and muon detector were designed and tested to observe how effective muon deflection is.

The implemented design shields from muons through a combination of natural metallic materials and a magnetic field produced by an arrangement of electromagnets and neodymium magnets. The metals implemented were chosen through a simulation which utilized the Bethe Equation (1) to parse through the periodic table and select elements with the characteristics required to obtain the desired change in energy. Based on practicality and financial feasibility, our design utilized a collection of aluminum and steel to reduce the muon energy (Figure 2). This setup only required the generation of a magnetic field under 1 Tesla which was obtainable with a current of 5-10 A and neodymium magnets (Condotti, 2013).

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ight
angle = rac{4\pi}{m_e c^2} \cdot rac{nz^2}{eta^2} \cdot \left(rac{c^2}{4\piarepsilon_0}
ight)^2 \cdot \left[\ln\left(rac{2m_e c^2 eta^2}{I \cdot (1-eta^2)}
ight) - eta^2
ight]$$

$$n = rac{N_A \cdot Z \cdot
ho}{A \cdot M_u}\,,$$

To assess the effectiveness of the muon shield, a small scale muon detector was constructed (Figure 6). The detector is composed of a PMT coupled with a scintillator, a charge sensitive amplifier, and a microcontroller wired to a laptop. A Photonis XP5312 PMT, supplied with 1000 V, detects photons emitted from a BC408 plastic scintillator and outputs a current. The amplifier (Figure 7) provides a gain of a factor of 10 and extends the pulse width to allow the Mega 2560 Microcontroller to record voltages at rate within its capabilities. The data from the serial monitor is processed through an analysis software written in Python.

The experimentation took place in the basement of Birge Hall at UC Berkeley. Experimentation was conducted over three trials. The first trial served as a control where the detector was completely uncovered. The second was under a metal barrier with a thickness of approximately 20-23 cm. The third was a combination of the metal barrier and magnetic field produced by neodymium magnets and electromagnets (Figure 1). The electromagnets were toroids wrapped with 24 gauge copper wire and powered by a 9 V supply. The neodymium magnets were positioned in a Halbach Array.



Figure 1: Experiment setup for the third trial. The **PMT is under the** neodymium magnets and electromagnet and the 20-23 cm metal barrier. **Photographed by Siddhant** Mehrotra, April 20th, 2018.

Designing an Electromagnetic Shield to Block Secondary Cosmic Rays

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The control trial (Figure 3) showed a regular and contained distribution centered at around 250 (~1.25 V) varying by about 50 (~0.25 V). Data points with a value above 1024 (~5 V), are outside the range of the input port and are rejected as a software error.

These datasets do not have any obvious spikes or other features that may indicate muon detections. The first trial of this set was strikingly different from the others, suggesting that it occurred under different circumstances. The data suggest a small amount of light was leaking into the PMT, causing the observed signal and masking any signal from muons.

The trial with only the metal shield (Figure 4) showed a dense distribution of points from 0 to 30 (0 to ~0.15 V) with some points up to about 80 (~0.4 V). However, after around point 40,000 (~200 s), the distribution changes abruptly. Most of the points after were between 0 and 5 (0 to ~0.025 V) with only a few up to a maximum of about 40 (~0.2 V). This sharp "cliff" was likely caused by a problem in the electronics. Again, the distribution did not have any features which would suggest muon detections.

The trial with both metal and magnetic shielding (Figure 5) initially showed a distribution that was dense between 0 to 25 (0 to ~ 0.125 V) with few points up to about 80 (~ 0.4 V). At around point 5,000 (~25 seconds), the distribution rises dramatically but smoothly. It settles at around 300 (~1.5 V) with very little variance except for a few vertical "spikes." These voltage spikes, extending upward and downward, could indicate muon detections. It should be noted that the data from the microcontroller with no input settles at around 325 (~1.5 V) with little variance, suggesting that this sudden rise might have again been an electronic failure and the spikes may not be meaningful.

Conclusion & Future Work

The objective of this experiment was to engineer a device for muon detection as well as a system for reducing the muon flux through an area using magnets and physical shields in order to facilitate the detection of other particles. After a design, testing, and revision process, a final system for detection and shielding was developed and assembled. During data collection and analysis, a python script that used numpy and matplotlib was used to analyze and plot the data. Due to potential problems with electronics, further experimentation is required to make a more accurate statement on the practicality of such a design. Future investigations would entail further debugging as well as an increased run time for each trial.

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detector

Figure 7: Charge sensitive preamplifier

+	N	
t(PMT)	+	Output(Micro-Controller)
10 κΩ		3 nF
	100 kΩ	<u> </u>
Î	100 kΩ	
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Analysis

Acknowledgments